Chapter 27.1

The Birth of Earth Video Worksheet

Solar System Formation

;;;;;;;;;;;;;;;;;;;;;;;;;;;;;;;;;;;;;;	(including the Earth) formed from a large, rotating cloud of interstellar 2.) Dust and 3.) called a 4.)
C	The in fall of material, increase in rotational speed and the crush of gravity created an enormous amount of 14.) Heat at the center. Ultimately, of 16.) Hydrogen into 17.) began, and
f t S ii Ii F S V t	formed our star, called the 18.)
F	Formation of Earth's Core
2 n	The Proto-Earth continued to grow from collisions, until the inner part of the protoplanet was hot enough to melt the 23.) Heavy elements, such as 24.) and 25.) Nickel. Due to their larger densities such (now liquid) netals began to sink to the Earth's center. This resulted in a separation of a 26.) only 10 million years

after the Earth began to form. This produced the layered structure of Earth and
also set up the formation of Earth's 28.) which protected and
still protects Earth's 29.) today, from being stripped away by the
Sun's solar winds.
During the accretion of material to the protoplanet, a cloud of gaseous 30.) Silico must have surrounded the Earth, to condense afterwards as solid rocks on the surface. What was left surrounding the planet was an early atmosphere of 31.) Light elements from the solar nebula, mainly hydrogen
and helium.
Formation of Earth's Moon
A rare characteristic of our planet is its large natural satellite, called the 32.) During the 33.) program, rocks from the Moon's surface were brought back to Earth. Radiometric dating of these rocks has shown the Moon to be 34.) billion years old, about 30 to 55 million years younger than other bodies in the solar system. The rocks were found to be very 35.), suggesting that they were heated to extremely high temperatures at one time. Another special feature is the relatively low 36.) of the Moon, which must mean it does not have a large metallic core, like all other terrestrial bodies in the solar system. In fact, the Moon has a bulk composition closely resembling the Earth's 37.) and crust together, without the Earth's core. This has led to the 38.), the idea that the Moon was formed during a giant impact of the 39.), the with another protoplanet. The Moon formed by accretion of the material blown off the mantles of the proto-Earth and impactor.
The impactor, sometimes named Theia, is thought to have been a little smaller than the current planet 40.) . The point at which it collided with Earth is an estimated 41.) 4.533 billion years ago. Models show that when an impactor this size struck the proto-Earth at a low angle and relatively low speed (8-20 km/sec), a lot of material from the mantles (and proto-crusts) of the proto-Earth and the impactor was ejected into space, where much of it stayed in orbit around the Earth. This material would eventually form the Moon. However, the metallic cores of the impactor would have sunk through the Earth's mantle to fuse with the Earth's core, depleting the Moon of metallic material. The giant impact hypothesis thus explains the Moon's abnormal 42.)

gigantic amount of energy, causing both the Earth and Moon to be completely molten. Immediately after the impact, the Earth's surface was a large 43.) Magma ocean. The impact is also thought to have changed Earth's 44.) to produce the large 23.5° tilt that is responsible for Earth's 45.)
Formation of Earth's Atmosphere and Oceans
The large impact would have blown off most of early atmosphere the Earth had. Because, the Earth lacked an atmosphere immediately after the giant impact, cooling would have been 46.) Fast. Within 150 million years a 47.) crust of lighter elements like Silica would have formed.
48.) Steam escaping from the crust, and other gases being released by 49.) completing the second atmosphere. Additional water was imported to Earth from numerous collisions, probably from asteroids ejected from the 50. Asteroid Belt under the influence of Jupiter's gravity.
All of the large amount of 51.) on Earth could never have been produced by volcanism and degassing alone. It is assumed that most of Earth's water was derived from millions of impacting 52.), which contain large amounts of 53.), and later collisions with small, icy protoplanets similar to today's icy moons. Impacts of these objects could have enriched the 54.) planets (Mercury, Venus, the Earth and Mars) with water, carbon dioxide, methane, ammonia, 55.) and other gases.
As the planet cooled, 56.) Clouds formed. 57.) gave rise to the oceans. Recent evidence suggests the oceans may have begun forming by 58.) 4.2 billion years ago. At the start of the Archaean eon, the Earth was already covered with 59.) The new atmosphere probably contained ammonia, methane, water vapor, carbon dioxide, and nitrogen, as well as smaller amounts of other gases. 60.) Oxygen was very scarce on the early Earth, until early 61.) ocean organisms evolved, like cyanobacteria in 62.) Standfiltes. Changing the Earth's atmosphere to an oxygen rich environment like we see today. This also formed an 63.) layer, blocking out harmful 64.) radiation, allowing life to move out of the oceans
and onto land.

Video Worksheet Word Bank

13.7	Magnetic Field
4.2	Mantle
4.527	Mars
4:555	Moon
4.54	Nickel_
4.6	Nitrogen
Apollo	Nuclear Fusion
Asteroid Belt	Oceans
Atmosphere	-Oxygen-
Axis	Ozone
-Big Bang	Photosynthetic
-Clouds-	Planetesimals
Comets	-Proto-Earth-
Composition	-Protoplanetary Disk -
Contract-	Protoplanets
Core	Protostar
Crust	Rain
Density	Seasons
-Dry-	Shock Wave
-Dust	- Silica
Earth	Solar Nebula
Fast	Solar System
Gas	Solar Wind
Giant Impact Hypothesis	Solid
-Heat	-Steam-
-Heavy-	-Stromatolites -
Helium	Sun
-Hydrogen_	Supernova
-Ice	Terrestrial
Iron	UV
Light	Volcanoes
-Magma-	Water